

## Solutions and Hints for the fourth problem set

**4.1. (30 points)**

**4.2. (30 points)**

The easiest way is to start from definition (eq. 5.57)

$$\theta = d(1+z)/D$$

where

$$D = R_c \sin(r/R_c)$$

for E-dS universe,  $R_c = \infty$ , so  $D = r$ , and  $r$  is the co-moving distance:

$$r = \int_{r_1}^{t_0} c dt / R(t)$$

Again, in an E-dS universe,  $R(t) \sim t^{2/3}$ , or  $t \sim R^{3/2} \sim (1+z)^{-3/2}$ . Plug those in, you can then show that:

$$r \sim 1 - (1+z)^{-1/2}$$

Then, find where  $\partial\theta/\partial z = 0$ .

**4.3. (40 point)**

This problem is asking you derive  $\Omega(z)$  for a  $\Lambda$  universe. Many of you assume  $\Lambda = 0$ , rather than  $\kappa = 0$ . The correct way is:

$$\Omega(z) = \frac{8\pi\rho(z)}{3H(z)^2} = \Omega_0 \frac{(1+z)^3}{\Omega_0(1+z)^3 + \Omega_\Lambda}$$

this uses the fact that  $\rho(z) = \rho_0(1+z)^3$ , and  $H(z)^2 = H_0^2[\Omega_0(1+z)^3 + \Omega_\kappa(1+z)^2 + \Omega_\Lambda]$ . and  $\Omega_\kappa = 0$ , and  $\Omega_\Lambda = 1 - 0.3$ . Then you will find that, at  $\Omega(z = 0.5) = 0.59$ ,  $\Omega(z = 10) = 0.998$ ,  $\Omega(R = 10) = 4 \times 10^{-4}$ .

**5.1. (25 point)**

Note that textbook seems to have an error. The correct answer should be  $\Omega_r = 2 \times 10^{-5} h^2$ .

**5.2. (25 point)**

**5.3. (25 point)**

**5.4. (25 point)** Integrate the Planck distribution, and you will find that  $n_r \sim 20T^3 \text{cm}^{-3}$ .